

StatMech pieces to the dark energy puzzle

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An old fine-tuning problem in high energy physics and cosmology concerns the very small but allegedly non-zero dark energy density Λ . In cosmic observations, the presence of Λ becomes noticeable — by causing an accelerated expansion of the universe — precisely at the epoch where the cosmic mass density becomes exceedingly stratified (the epoch of “structure formation”). Consider now that...

- 1) the aforementioned observed cosmic inhomogeneities are mostly neglected in the standard “ Λ CDM” model of cosmology.
- 2) the value of Λ is of the same order of magnitude as the mass density of the visible “baryonic” mass. That prompts the following question: could a statistical mechanical treatment of cosmic inhomogeneity possibly shed some light on the dark energy problem?

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