

# Universal features of Lifshitz Green's functions from holography

Gino Knodel

University of Michigan

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C. Keeler, G.K., J.T. Liu, and K. Sun; arXiv:1505.07830.

C. Keeler, G.K., and J.T. Liu; arXiv:1404.4877.

# Motivation

Challenges of AdS/CMT:

- ▶ Holographic dictionary less clear than in relativistic case.
- ▶ Holographic models often describe an idealized theory.

→ Universal predictions of nonrelativistic gauge/gravity duality (e.g. analog of  $\eta/s$ )?

- ▶ Main focus: Holographic Green's functions.
  1. What do we know?
  2. Which features are robust w.r.t. deformations of the theory?
  3. Comparison with field theory

## Constraining holographic observables via symmetries

- ▶ Type IIB on  $AdS_5 \times S^5 \Leftrightarrow \mathcal{N} = 4$  SYM.

Consider retarded Green's function of scalar operator  $\mathcal{O}_\Delta$ :

$$G_R(\omega, \vec{k}) = -i \int d^{d+2}x e^{-ik \cdot x} \theta(t) \langle [\mathcal{O}_\Delta(x), \mathcal{O}_\Delta(0)] \rangle$$

Superconformal symmetry  $\implies G_R(q^2) = A(-q^2)^{\Delta-2},$   
 $q^2 = \omega^2 - |\vec{k}|^2, \quad A = \text{const.}$

- ▶ What about non-relativistic symmetries?
  - ▶ Lifshitz scaling

## Lifshitz Green's functions

Lifshitz scaling symmetry:

$$\vec{x} \rightarrow \Lambda^{\frac{1}{z}} \vec{x}, \quad t \rightarrow \Lambda t$$

$z > 1$ : dynamical critical exponent

Assuming rotational invariance:

$$G_R(\omega, \vec{k}) = |\vec{k}|^{2\nu z} \mathcal{G}(\hat{\omega}), \quad \hat{\omega} = \frac{\omega}{|\vec{k}|^z}$$

Properties of Green's functions in Lifshitz field theories without additional symmetries?

## Lifshitz Green's functions - general properties

$$G_R(\omega, \vec{k}) = |\vec{k}|^{2\nu z} \mathcal{G}(\hat{\omega})$$

Generic features:

1.  $\mathcal{G}$  is analytic in the upper half plane (causality)
2. Scaling behavior:
  - 2.1  $\mathcal{G}(\hat{\omega} \rightarrow 0) \sim \text{const.}$  (so that  $G_R \sim |\vec{k}|^{2\nu z}$  as  $\omega \rightarrow 0$ )
  - 2.2  $\mathcal{G}(\hat{\omega} \rightarrow \infty) \sim \hat{\omega}^{2\nu}$  (so that  $G_R \sim \omega^{2\nu}$  as  $|\vec{k}| \rightarrow 0$ )

No (obvious) further constraints.

“Lifshitz field theory”: Family of theories with different dynamics, and thus  $\mathcal{G}(\hat{\omega})$ .

## Holographic Lifshitz Green's function

However: On the gravity side, there is a “canonical” result.

Lifshitz spacetime:

$$ds_{d+2}^2 = \frac{-dt^2 + d\rho^2}{\rho^2} + \frac{d\vec{x}^2}{\rho^{2/z}}$$

$$\text{isometry: } \vec{x} \rightarrow \Lambda^{\frac{1}{z}} \vec{x}, \quad t \rightarrow \Lambda t, \quad \rho \rightarrow \Lambda \rho.$$

For  $z = 2$ :

$$\mathcal{G}(\hat{\omega}) = K_0 \hat{\omega}^{2\nu} \frac{\Gamma\left(\frac{1}{2} + \nu + \frac{i}{4\hat{\omega}}\right)}{\Gamma\left(\frac{1}{2} - \nu + \frac{i}{4\hat{\omega}}\right)}$$

[Kachru, Liu, Mulligan; 0808.1725]

Where is the freedom on the gravity side?

## Review: The two-derivative holographic Green's function

Consider a probe scalar in Lifshitz spacetime:

$$(\square - m^2)\phi = 0$$

Ansatz:  $\phi = e^{i(\vec{k}\cdot\vec{x} - \omega t)} \rho^{d/2z} \psi(\rho) \implies$  Schrödinger-like equation:

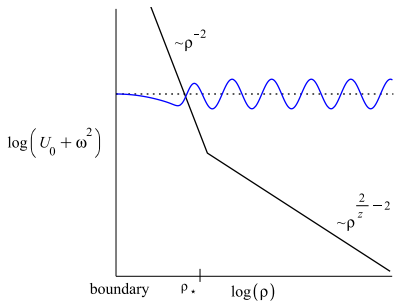
$$-\psi''(\rho) + U_0\psi(\rho) = 0$$

$$U_0 = \frac{\nu^2 - \frac{1}{4}}{\rho^2} + \frac{|\vec{k}|^2}{\rho^{2-2/z}} - \omega^2, \quad \nu^2 = m^2 + \left(\frac{d+z}{2z}\right)^2$$

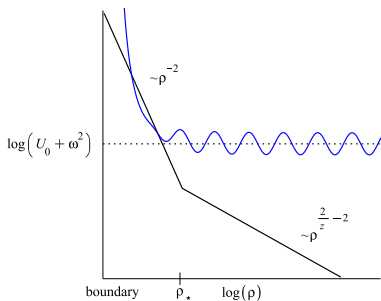
Using scale-invariant coordinates  $\hat{\rho} = \rho|\vec{k}|^z$ :

$$-\psi''(\hat{\rho}) + \hat{U}_0(\hat{\rho})\psi(\hat{\rho}) = 0, \quad \hat{U}_0(\hat{\rho}) = \frac{\nu^2 - \frac{1}{4}}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2$$

normalizable mode:



non-normalizable mode:



Near the horizon:

$$\psi(\hat{\rho} \rightarrow \infty) \sim ae^{i\hat{\omega}\hat{\rho}} + be^{-i\hat{\omega}\hat{\rho}}$$

Near the boundary:

$$\psi(\hat{\rho} \rightarrow 0) \sim A\hat{\rho}^{\frac{1}{2}-\nu} + B\hat{\rho}^{\frac{1}{2}+\nu}$$



## Holographic Green's functions

$$\psi(\hat{\rho} \rightarrow \infty) \sim ae^{i\hat{\omega}\hat{\rho}} + be^{-i\hat{\omega}\hat{\rho}}$$

$$\psi(\hat{\rho} \rightarrow 0) \sim A\hat{\rho}^{\frac{1}{2}-\nu} + B\hat{\rho}^{\frac{1}{2}+\nu}$$

Choose infalling boundary conditions for retarded Green's function:

$$\mathcal{G}(\hat{\omega}) = \hat{\omega}^{2\nu} \frac{B}{A} \Big|_{b=0}$$

Find relation between  $\{A, B\}$  and  $\{a, b\}$  by solving a quantum mechanical scattering problem.

## Introducing higher derivatives

Where does the freedom of finding different  $\mathcal{G}$  arise in the gravity theory?

1. Choice of background metric?
2. Dynamics of probe scalar, e.g. **higher derivative corrections**:

$$\blacktriangleright \left(\frac{\partial}{\partial t}\right)^i \hat{=} \omega^i, \quad |\vec{\nabla}|^j \hat{=} |\vec{k}|^j:$$

modified Schrödinger potential

$$\begin{aligned}\hat{U} &= \hat{U}_0(\hat{\rho}) + \frac{1}{\hat{\rho}^2} \sum_{i+j>2} \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z}, & \hat{\rho} &= \rho |\vec{k}|^z, \quad \hat{\omega} = \frac{\omega}{|\vec{k}|^z} \\ &= \hat{U}_0(\hat{\rho}) + \frac{1}{\hat{\rho}^2} \sum_{i+j>2} \lambda_{i,j} (\omega \rho)^i (|\vec{k}| \rho^{1/z})^j\end{aligned}$$

where  $\lambda_{i,j} \sim \left(\frac{\ell}{L}\right)^{i+j-2}$ ,  $\ell \ll L$  (curvature scale).

$$\blacktriangleright \left(\frac{\partial}{\partial \rho}\right)^n:$$

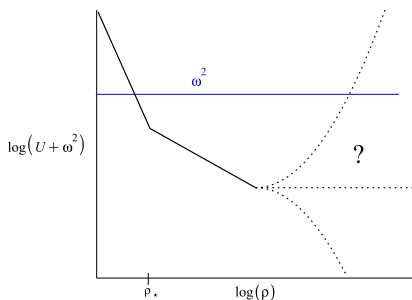
$$\begin{aligned}\text{e.g. :} \quad & -\psi'' + \hat{U}\psi = \lambda\psi^{(4)} \\ \Rightarrow & -\psi'' + U\psi - \lambda(\hat{U}\psi)'' = \mathcal{O}(\lambda^2) \\ \Rightarrow & -\tilde{\psi}'' + \tilde{U}\tilde{\psi} \approx 0\end{aligned}$$

Higher derivative corrections can be captured by modifying the effective potential.

## Keeping higher derivatives under control

Consider only a single higher derivative term ( $i$  temporal,  $j$  spatial derivatives):

$$\hat{U}(\hat{\rho}) = \hat{U}_0(\hat{\rho}) + \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z-2}$$



Want to compute Green's function perturbatively, i.e.

$$\mathcal{G} = \mathcal{G}_0 + \delta\mathcal{G}.$$

Problem: higher derivatives generically blow up near horizon ( $\hat{\rho} \rightarrow \infty$ ).

But: Green's function is given by  $B/A$  (near-boundary behavior of  $\psi$ ) and thus only contains information about the effective potential **in the tunneling regime**.

## WKB approximation

Recall:

$$\mathcal{G}(\hat{\omega}) = \hat{\omega}^{2\nu} \frac{B}{A} \Big|_{b=0} \quad \psi(\hat{\rho} \rightarrow 0) \sim A\hat{\rho}^{\frac{1}{2}-\nu} + B\hat{\rho}^{\frac{1}{2}+\nu}$$

Infalling boundary conditions can be imposed order by order in  $\lambda_{i,j}$ .

WKB approximation :

$$\psi_{WKB}(\hat{\rho}) = \begin{cases} \sqrt{\nu} \hat{U}^{-\frac{1}{4}} \left( C_1(a) e^{S(\hat{\rho}, \hat{\rho}_0)} + C_2(a) e^{-S(\hat{\rho}, \hat{\rho}_0)} \right) & , \hat{\rho} < \hat{\rho}_0 \\ a(-\hat{U})^{-\frac{1}{4}} e^{i\Phi(\hat{\rho}, \hat{\rho}_0)} & , \hat{\rho} > \hat{\rho}_0 \end{cases}$$

$$S(\hat{\rho}, \hat{\rho}_0) = \int_{\hat{\rho}}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\hat{U}}, \quad \Phi(\hat{\rho}, \hat{\rho}_0) = \int_{\hat{\rho}_0}^{\hat{\rho}} d\hat{\rho} \sqrt{-\hat{U}}$$

$\hat{\rho}_0$  : classical turning point.

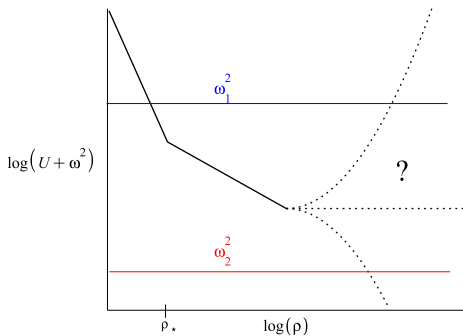
$$\psi_{WKB}(\hat{\rho}) = \begin{cases} \sqrt{\nu} \hat{U}^{-\frac{1}{4}} \left( C_1(a) e^{S(\hat{\rho}, \hat{\rho}_0)} + C_2(a) e^{-S(\hat{\rho}, \hat{\rho}_0)} \right) & , \hat{\rho} < \hat{\rho}_0 \\ a(-\hat{U})^{-\frac{1}{4}} e^{i\Phi(\hat{\rho}, \hat{\rho}_0)} & , \hat{\rho} > \hat{\rho}_0 \end{cases}$$

Coefficients of non-normalizable/normalizable mode:

$$A = \lim_{\epsilon \rightarrow 0} C_1 \epsilon^\nu e^{S(\epsilon, \rho_0)}, \quad B = \lim_{\epsilon \rightarrow 0} C_2 \epsilon^{-\nu} e^{-S(\epsilon, \rho_0)}$$

$$\begin{aligned} S(\epsilon, \hat{\rho}_0) &= \int_{\epsilon}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\hat{U}} \\ &= \int_{\epsilon}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\frac{\nu^2}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2 + \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z-2}} \end{aligned}$$

- ▶ Coefficients  $A, B$  contain information about  $\hat{U}$  up to the classical turning point  $\hat{\rho}_0$ .
  - ▶ Can calculate Green's function with higher derivatives, as long as h.d. are subdominant compared to two-derivative terms.



$$S = \int_{\epsilon}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\frac{\nu^2}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2 + \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z-2}}$$

$$\stackrel{!}{=} S_0 + \delta S + O(\lambda_{i,j}^2)$$

Will have to impose upper bound on  $\lambda_{i,j}$  **and** lower cutoff for  $\hat{\omega}$ .

## Calculating the spectral function

Calculate  $\text{Im}\mathcal{G}$  (spectral function) via

$$\text{Im}\mathcal{G}(\hat{\omega}) = \hat{\omega}^{2\nu} \text{Im} \left( \frac{B}{A} \right) = K_0 \hat{\omega}^{2\nu} \lim_{\epsilon \rightarrow 0} \epsilon^{-2\nu} e^{-2S(\epsilon, \hat{\rho}_0)}$$

(can find  $\text{Re}\mathcal{G}$  via Kramers-Kronig relation)

Can expand

$$\text{Im}\mathcal{G} = \text{Im}\mathcal{G}_0 + \delta\text{Im}\mathcal{G} + \mathcal{O}(\lambda^2)$$

by expanding integral

$$S(\hat{\rho}, \hat{\rho}_0) = \int_{\hat{\rho}}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\hat{U}_0 + \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z-2}}$$

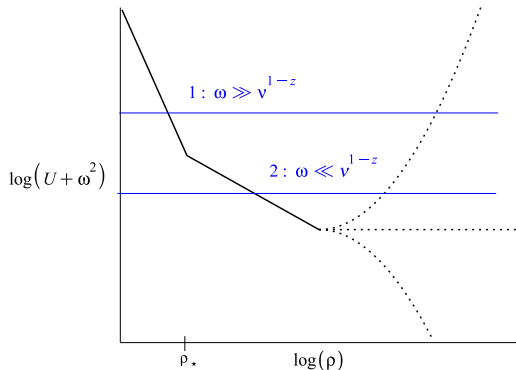
linearly in  $\lambda_{i,j}$ .

## Results

Recall:

$$\hat{U}(\hat{\rho}) = \frac{\nu^2 - \frac{1}{4}}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2 + \lambda_{i,j} \hat{\omega}^j \hat{\rho}^{i+j/z-2}$$

Two interesting limits:





## High frequency limit - two derivative result

1.  $\hat{\omega} \gg \nu^{1-z}$ :

Classical turning point lies at  $\hat{\rho}_0 \sim \frac{\nu}{\hat{\omega}}$ .

To leading order:

$$S_0 = \int_{\epsilon}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\frac{\nu^2}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2} \approx \int_{\epsilon}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\frac{\nu^2}{\hat{\rho}^2} - \hat{\omega}^2} \sim -\nu \log(\epsilon) + \text{const.}$$

Gives rise to power law scaling near the boundary:

$$\begin{aligned} \psi(\epsilon \ll 1) &\approx \epsilon^{\frac{1}{2}} \left( C_1 e^{S_0(\epsilon, \hat{\rho}_0)} + C_2 e^{-S_0(\epsilon, \hat{\rho}_0)} \right) \\ &\sim \tilde{C}_1 \epsilon^{\frac{1}{2}-\nu} + \tilde{C}_2 \epsilon^{\frac{1}{2}+\nu} \end{aligned}$$

2-derivative Green's function:

$$\text{Im} \mathcal{G}(\hat{\omega}) = \hat{\omega}^{2\nu} \text{Im} \left( \frac{B}{A} \right) = \text{Im} \left( \frac{\tilde{C}_2}{\tilde{C}_1} \right) \hat{\omega}^{2\nu}$$

## High frequency limit with higher derivatives

Computing  $S$  to linear order in  $\lambda$  yields:

$$\text{Im}\mathcal{G}(\hat{\omega}) \approx C\hat{\omega}^{2\nu}, \quad C = (2\nu)^{-2\nu} \exp \left[ 2\nu(1 - \delta_{j,0} c_i \lambda_{i,0} \nu^{j-2} + \dots) \right]$$

- ▶ Expected large frequency behavior  $G_R = |\vec{k}|^{2\nu} \mathcal{G}(\hat{\omega}) \sim \omega^{2\nu}$  ("featureless"; similar to AdS)
  - ▶ Higher derivatives normalize prefactor. Spatial derivatives derivatives ( $j \neq 0$ ) are subleading in the  $\hat{\omega} \rightarrow \infty$  limit.
  - ▶ H.d. corrections become of order one if

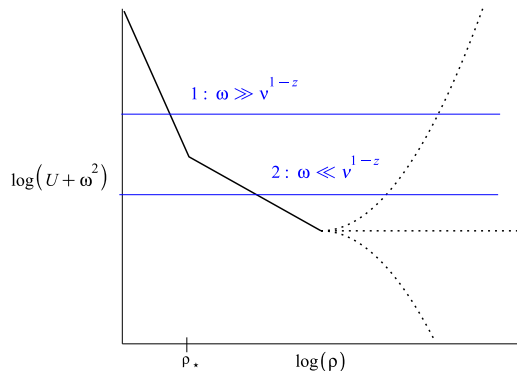
$$\lambda_{i,0} \nu^{j-2} \sim 1 \quad \Leftrightarrow \quad \delta U_{hd}(\hat{\rho}_0) = \lambda_{i,0} \hat{\omega}^i \hat{\rho}_0^{i-2} \sim \hat{\omega}^2$$

- ▶ Size of corrections to spectral function is controlled by relative size of h.d. corrections to the potential up to the classical turning point.
- ▶ Translates into condition on mass:  $\lambda_{i,0} \nu^{j-2} \sim (\frac{\ell}{L} \nu)^{i-2} \sim (m\ell)^{i-2} \ll 1$

## Low frequency limit

Effective potential

$$\hat{U}(\hat{\rho}) = \frac{\nu^2 - \frac{1}{4}}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2 + \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z-2}$$



## Low frequency limit - two derivative result

$$2. \hat{\omega} \ll \nu^{1-z};$$

Classical turning point lies at  $\hat{\rho}_0 \sim \hat{\omega}^{-\frac{z}{z-1}}$ .

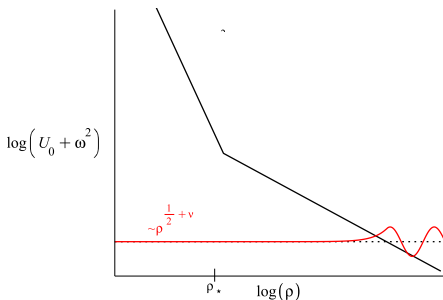
To leading order:

$$\begin{aligned} S_0 &= \int_{\epsilon}^{\hat{\rho}_0} d\hat{\rho} \sqrt{\frac{\nu^2}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2} \sim -\nu \log(\epsilon) + \int^{\hat{\omega}^{-\frac{z}{z-1}}} d\hat{\rho} \frac{1}{\hat{\rho}^{1-1/z}} + \text{const.} \\ &\sim -\nu \log(\epsilon) + \hat{\omega}^{-\frac{1}{z-1}} + \text{const.} \end{aligned}$$

Wavefunction near the boundary:

$$\begin{aligned} \psi(\epsilon \ll 1) &\approx \epsilon^{\frac{1}{2}} \left( C_1 e^{S_0(\epsilon, \hat{\rho}_0)} + C_2 e^{-S_0(\epsilon, \hat{\rho}_0)} \right) \\ &\sim \tilde{C}_1 \epsilon^{\frac{1}{2}-\nu} \exp\left(\frac{E_0}{2} \hat{\omega}^{-\frac{1}{z-1}}\right) + \tilde{C}_2 \epsilon^{\frac{1}{2}+\nu} \exp\left(-\frac{E_0}{2} \hat{\omega}^{-\frac{1}{z-1}}\right) \end{aligned}$$

e.g. normalizable mode:



$$\psi(\epsilon \ll 1) \sim \tilde{C}_1 \epsilon^{\frac{1}{2} - \nu} \exp\left(\frac{E_0}{2} \hat{\omega}^{-\frac{1}{z-1}}\right) + \tilde{C}_2 \epsilon^{\frac{1}{2} + \nu} \exp\left(-\frac{E_0}{2} \hat{\omega}^{-\frac{1}{z-1}}\right)$$

→ exponential growth/decay due to tunneling under  $\sim \hat{\rho}^{2/z-2}$  barrier.

2-derivative Green's function:

$$\text{Im}\mathcal{G}(\hat{\omega}) = D \hat{\omega}^{2\nu} \exp[-E_0 \hat{\omega}^{-\frac{1}{z-1}}].$$

## Low frequency limit with higher derivatives

Computing  $S$  to linear order in  $\lambda$  yields:

$$\text{Im}\mathcal{G}(\hat{\omega}) = D\hat{\omega}^{2\nu} \exp[-\hat{\omega}^{-\frac{1}{z-1}}(E_0 + \delta E(\hat{\omega}))]$$

$$D = (2\nu)^{-2z\nu} z^{2\nu(1-z)} \exp\left[2z\nu(1 + \delta_{i,0} d_j \lambda_{0,j} \nu^{j-2} + \dots)\right], \quad d_j = \int_{x=0}^{\infty} dx \frac{x^{\frac{1}{2}-1}}{\sqrt{1+x}}$$

$$E_0 = \frac{\sqrt{\pi}\Gamma\left(\frac{1}{2(z-1)}\right)}{z\Gamma\left(\frac{z}{2(z-1)}\right)}, \quad \delta E(\hat{\omega}) = e_{i,j} \lambda_{i,j} \hat{\omega}^{-\frac{1}{z-1}(i+j-2)} + \dots, \quad e_{i,j} = \int_0^1 dx \frac{x^{i-1+\frac{j-1}{z}}}{\sqrt{1-x^{2-\frac{2}{z}}}}$$

- ▶ Prefactor is renormalized by terms  $\sim \lambda_{0,j} \nu^{j-2}$ .
- ▶ Exponent is modified by frequency-dependent terms. Recall:

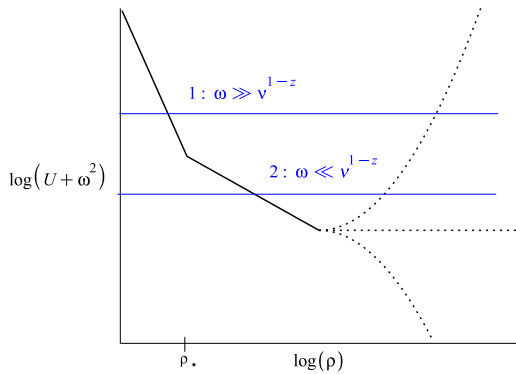
$$\hat{U}(\hat{\rho}) = \frac{\nu^2 - \frac{1}{4}}{\hat{\rho}^2} + \frac{1}{\hat{\rho}^{2-2/z}} - \hat{\omega}^2 + \lambda_{i,j} \hat{\omega}^i \hat{\rho}^{i+j/z-2}$$

Classical turning point lies at

$$\hat{\rho}_0 \approx \hat{\omega}^{-\frac{z}{z-1}}$$

H.d. corrections become of order one if

$$\lambda_{i,j} \hat{\omega}^{-\frac{1}{z-1}(i+j-2)} \sim 1 \quad \Leftrightarrow \quad \lambda_{i,j} \hat{\omega}^i \hat{\rho}_0^{i+j/z-2} \sim \hat{\omega}^2$$



## Exponential suppression

At low frequencies:

$$\text{Im}G(\hat{\omega} \ll \nu^{1-z}) \approx D\hat{\omega}^{2\nu} \exp[-\hat{\omega}^{-\frac{1}{z-1}} (E_0 + \delta E(\hat{\omega}))]$$

$$\delta E(\hat{\omega}) = e_{i,j} \lambda_{i,j} \hat{\omega}^{-\frac{1}{z-1}(i+j-2)} + \dots$$

Strict low-frequency limit  $\hat{\omega} \rightarrow 0$  is non-universal (h.d. dominate), but corrections are under control if

$$\lambda_{i,j} \hat{\omega}^{-\frac{1}{z-1}(i+j-2)} \ll 1.$$

Using  $\lambda_{i,j} \sim \left(\frac{\ell}{L}\right)^{i+j-2}$ :

$$\nu^{1-z} \gg \hat{\omega} \gg \left(\frac{\ell}{L}\right)^{z-1}$$

- ▶ Recall:  $\nu \ll \frac{L}{\ell} \Leftrightarrow m \ll \frac{1}{\ell} \rightarrow$  exponential suppression is a robust feature in a wide window of frequencies.
- ▶ Corrections can be computed order-by-order for different models.



## Interpretation of exponential behavior

- ▶ Exponential behavior of  $\text{Im}\mathcal{G}$  is a robust result at low frequencies.
- ▶ Can we find the same behavior of  $\text{Im}\mathcal{G}$  on the field theory side?

## Example: Quadratic band crossing model

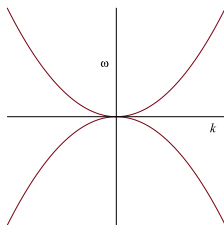
Scale-invariant theory with  $z = 2$ :

$$S = \int d^2x dt \left\{ \bar{\Psi} [i\gamma_0 \partial_0 + \gamma_1 (\partial_x^2 - \partial_y^2) + 2\gamma_2 \partial_x \partial_y] \Psi - g \psi_1^\dagger \psi_2^\dagger \psi_2 \psi_1 \right\}$$

where  $\Psi = (\psi_1, \psi_2)^T$ .

Dispersion relation:

$$\epsilon_{\pm}(\vec{k}) = \pm |\vec{k}|^2$$



Can consider bosonic particle-hole operators

$$b_i = \bar{\Psi} \gamma_i \Psi, \quad i = 0, 1, 2, 3$$

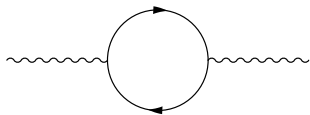
## Calculating self-energy corrections

Want to calculate spectral weight  $\text{Im}G_{(b)}$ .

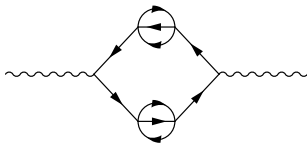
Optical theorem:

$$\text{Im}G_{(b)}(\omega, \vec{k}) \neq 0 \quad \Leftrightarrow \quad \Gamma_{b_i \rightarrow \dots}(\omega, \vec{k}) \neq 0$$

1-loop:



5-loop:



etc.

At n-loop order,  $O(g^{2n})$ ,  $\text{Im}G_{(b)}$  is related to decay rate for

$$b\{\omega, \vec{k}\} \rightarrow n \text{ particles}\{\omega_{p_i}, \vec{k}_{p_i}\} + n \text{ holes}\{\omega_{h_i}, \vec{k}_{h_i}\}.$$

Energy and momentum conservation:

$$\vec{k} = \sum_{i=1}^{n+1} \vec{k}_{p_i} - \sum_{i=1}^{n+1} \vec{k}_{h_i},$$
$$\omega = \sum_{i=1}^{n+1} \omega_{p_i} - \sum_{i=1}^{n+1} \omega_{h_i} = \sum_{i=1}^{n+1} (|\vec{k}_{p_i}|^2 + |\vec{k}_{h_i}|^2) \geq \frac{k^2}{2(n+1)},$$

At each fixed order, spectral weight is zero below a certain threshold.

→ for fixed  $\omega \ll \vec{k}^2$ , spectral weight can only arise via process of  $O(g^{2n})$ , where  $n \sim \frac{k^2}{2\omega}$

$$\Rightarrow \text{Im}G_{(b)}(\omega \ll \vec{k}) \sim g^{2n} \sim g^{k^2/\omega}$$

For  $g \ll 1$  this implies

$$\text{Im}G_{(b)}(\omega \ll \vec{k}^2) \sim \exp\left(-\frac{\text{const.}}{\hat{\omega}}\right), \quad \hat{\omega} = \frac{\omega}{|\vec{k}|^2}.$$

Result can be generalized to  $z \neq 2$ :

$$\text{Im}G_{(b)}(\omega \ll \vec{k}^2) \sim g^{\hat{\omega}^{-\frac{1}{z-1}}} \sim \exp(-\text{const.} \cdot \hat{\omega}^{-\frac{1}{z-1}}), \quad \hat{\omega} = \frac{\omega}{|\vec{k}|^z}.$$

- ▶ Agrees with the holographic prediction  $\text{Im}\mathcal{G}(\hat{\omega}) \sim \exp[-E_0 \hat{\omega}^{-\frac{1}{z-1}}]$ .
- ▶ Exp. behavior is generic for Lifshitz theories with bosonic decay channels.
- ▶ Aside: In Dirac theory ( $z = 1$ ), dispersion relation is  $\omega = |\vec{k}|$ , so energy conservation implies

$$\begin{aligned} \vec{k} &= \sum_{i=1}^{n+1} \vec{k}_{p_i} - \sum_{i=1}^{n+1} \vec{k}_{h_i}, \\ \omega &= \sum_{i=1}^{n+1} \omega_{p_i} - \sum_{i=1}^{n+1} \omega_{h_i} = \sum_{i=1}^{n+1} (|\vec{k}_{p_i}| + |\vec{k}_{h_i}|) \geq |\vec{k}|, \\ &\implies \text{Im}G_{(b)}(\omega < |\vec{k}|) = 0 \end{aligned}$$

## Summary

- ▶ Field theory: Family of theories with Lifshitz symmetry.
- ▶ Gravity: Family of theories with Lifshitz symmetry; dynamical information encoded in higher derivative couplings  $\lambda_{i,j}$ .
- ▶ Higher derivative corrections can be computed systematically:
  - ▶  $\text{Im}\mathcal{G}(\hat{\omega} \gg \nu^{1-z}) \sim C\hat{\omega}^{2\nu}$ ;
    - ▶  $C$  has perturbative expansion in  $\lambda\nu^{i-2} \sim (m\ell)^{i-2} \ll 1$ .
  - ▶  $\text{Im}\mathcal{G}(\hat{\omega} \ll \nu^{1-z}) \sim D\hat{\omega}^{2\nu} \exp[-E\hat{\omega}^{-\frac{1}{z-1}}]$ 
    - ▶  $D$  has perturbative expansion in  $\lambda\nu^{j-2}$ .
    - ▶  $E$  has perturbative expansion in  $\lambda_{i,j}\hat{\omega}^{-\frac{1}{z-1}(i+j-2)}$  ..
- ▶ “Naive” limit  $\hat{\omega} \rightarrow 0$  is not under perturbative control, but results are robust after imposing the cutoff  $\hat{\omega} \gg \left(\frac{\ell}{L}\right)^{z-1}$ .
- ▶ Can confirm exponential behavior of spectral function in a broad class of field theories.

## Open questions

1. Constraints on sign of  $\lambda_{i,j}$  from bulk causality/unitarity?
2. Compute corrections to  $\text{Im}\mathcal{G}(\hat{\omega})$  in string embeddings?
3. Measure spectral function experimentally?

Thank you!